

2. A basis of  $\ker(A^T)$  is  $\begin{bmatrix} 1 \\ -2 \\ 1 \end{bmatrix}$ .  $\text{im}(A)$  is the plane perpendicular to this line.

10. a. If  $\vec{x}$  is an arbitrary solution of the system  $A\vec{x} = \vec{b}$ , let  $\vec{x}_h = \text{proj}_V \vec{x}$ , where  $V = \ker(A)$ , and  $\vec{x}_0 = \vec{x} - \text{proj}_V \vec{x}$ . Note that  $\vec{b} = A\vec{x} = A(\vec{x}_h + \vec{x}_0) = A\vec{x}_h + A\vec{x}_0 = A\vec{x}_0$ , since  $\vec{x}_h$  is in  $\ker(A)$ .

b. If  $\vec{x}_0$  and  $\vec{x}_1$  are two solutions of the system  $A\vec{x} = \vec{b}$ , both from  $(\ker A)^\perp$ , then  $\vec{x}_1 - \vec{x}_0$  is in the subspace  $(\ker A)^\perp$  as well. Also,  $A(\vec{x}_1 - \vec{x}_0) = A\vec{x}_1 - A\vec{x}_0 = \vec{b} - \vec{b} = \vec{0}$ , so that  $\vec{x}_1 - \vec{x}_0$  is in  $\ker(A)$ . By Fact 5.1.8b, it follows that  $\vec{x}_1 - \vec{x}_0 = \vec{0}$ , or  $\vec{x}_1 = \vec{x}_0$ , as claimed.

c. Write  $\vec{x}_1 = \vec{x}_h + \vec{x}_0$  as in part a; note that  $\vec{x}_h$  is orthogonal to  $\vec{x}_0$ . The claim now follows from the Pythagorean Theorem (Fact 5.1.9).

24. Using Fact 5.4.6, we find  $\vec{x}^* = [2]$ .

34. We want  $\begin{bmatrix} c_0 \\ c_1 \\ c_2 \\ c_3 \\ c_4 \end{bmatrix}$  such that

$$\begin{bmatrix} 1 & 0 & 1 & 0 & 1 \\ 1 & \sin(0.5) & \cos(0.5) & \sin(1) & \cos(1) \\ 1 & \sin(1) & \cos(1) & \sin(2) & \cos(2) \\ 1 & \sin(1.5) & \cos(1.5) & \sin(3) & \cos(3) \\ 1 & \sin(2) & \cos(2) & \sin(4) & \cos(4) \\ 1 & \sin(2.5) & \cos(2.5) & \sin(5) & \cos(5) \\ 1 & \sin(3) & \cos(3) & \sin(6) & \cos(6) \end{bmatrix} \begin{bmatrix} c_0 \\ c_1 \\ c_2 \\ c_3 \\ c_4 \end{bmatrix} = \begin{bmatrix} 0 \\ 0.5 \\ 1 \\ 1.5 \\ 2 \\ 2.5 \\ 3 \end{bmatrix}$$

Since the columns of the coefficient matrix are linearly independent, its kernel is  $\{\vec{0}\}$ .

We can use Fact 5.4.6 to compute  $\begin{bmatrix} c_0 \\ c_1 \\ c_2 \\ c_3 \\ c_4 \end{bmatrix} \approx \begin{bmatrix} 1.5 \\ 0.109 \\ -1.537 \\ 0.303 \\ 0.043 \end{bmatrix}$  so  $f^*(t) \approx 1.5 + 0.109 \sin(t) - 1.537 \cos(t) + 0.303 \sin(2t) + 0.043 \cos(2t)$ .

40. First we look for  $\begin{bmatrix} c_0 \\ c_1 \end{bmatrix}$  such that  $\log D = c_0 + c_1 \log a$ .

Proceeding as in Exercise 39, we get  $\begin{bmatrix} c_0 \\ c_1 \end{bmatrix}^* \approx \begin{bmatrix} 0 \\ 1.5 \end{bmatrix}$ , i.e.  $\log D \approx 1.5 \log a$ , hence  $D \approx 10^{1.5 \log a} = a^{1.5}$ .

Note that the formula  $D = a^{1.5}$  is Kepler's third law of planetary motion.

